

Theory of Superconductivity, Frühjahrssemester 2026

Blatt 1: Self-study

The purpose of problems 1 and 2 is to review the basic properties of the ideal Fermi and Bose gases. This is a reading assignment; use any textbook or other source.

(1) Free (= ideal) Fermi gas

Use a textbook of your choice (e.g., S.H. Simon, The Oxford solid state basics, pages 27 – 32) to review the basic properties of the free Fermi gas. In particular, you should understand and remember

(a) the meaning and the shape of the Fermi function $f(E) = \frac{1}{\exp((E - \mu)/k_B T) + 1}$.

(b) the meaning and typical numerical values of the Fermi energy ϵ_F , the Fermi wavevector k_F , the Fermi velocity v_F , the Fermi temperature $T_F = \epsilon_F/k_B$.

(c) the relation $k_F^3 = 3\pi^2 n$ between the Fermi wavevector and the density $n = N/V$ of the fermions.

(d) the meaning of the density of states (per unit volume; for one spin direction)

$$N(E) = \frac{3}{4} \frac{n}{\epsilon_F} \left(\frac{E}{\epsilon_F} \right)^{1/2} \quad \text{and its value at the Fermi energy } N(\epsilon_F) = \frac{mk_F}{2\pi^2 \hbar^2}.$$

(e) the specific heat C at low temperatures $T \ll T_F$, $C = \gamma T$ where $\gamma = \frac{2\pi^2}{3} k_B^2 N(\epsilon_F)$.

(f) Pauli (spin) paramagnetism of the free electron gas, $\chi_{\text{Pauli}} = \mu_0 \frac{dM}{dB} = 2\mu_0 \mu_B^2 N(\epsilon_F)$ where $\mu_B = |e|\hbar/(2m)$ is the Bohr magneton. The orbital motion of the electrons leads to Landau diamagnetism, $\chi_{\text{Landau}} = -\frac{1}{3}\chi_{\text{Pauli}}$. The typical numerical values of χ_{Pauli} and χ_{Landau} are of order 10^{-6} .

(2) Free (= ideal) Bose gas

Use a textbook of your choice (e.g., Tilley², Superfluidity and Superconductivity, pages 31 – 38) to review the basic properties of the free Bose gas. In particular, you should understand and remember

(a) the meaning and the shape of the Bose function $f_B(E) = \frac{1}{\exp((E - \mu)/k_B T) - 1}$.

(b) Bose-Einstein condensation: the lowest one-particle state will be occupied by a macroscopic number $N_0(T)$ of particles at temperatures

$$T < T_B = \frac{2\pi\hbar^2}{mk_B} \left(\frac{N}{\zeta(3/2)V} \right)^{2/3} = 3.31 \frac{\hbar^2}{mk_B} n^{2/3}. \quad \text{In fact, } \frac{N_0(T)}{N} = 1 - \left(\frac{T}{T_B} \right)^{3/2}.$$